## Measurement of Neutral-Current $K^+$ Production by Neutrinos using MINERvA

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Neutral-current production of  $K^+$  by atmospheric neutrinos is a background in searches for the proton decay  $p \to K^+ \bar{\nu}$ . Reactions such as  $\nu p \to \nu K^+ \Lambda$  are indistinguishable from proton decays when the decay products of the  $\Lambda$  are below detection threshold. Events with  $K^+$  are identified in MINERvA by reconstructing the timing signature of a K<sup>+</sup> decay at rest. A sample of 201 neutrino-induced neutral-current  $K^+$  events is used to measure differential cross sections with respect to the  $K^+$  kinetic energy, and the non- $K^+$  hadronic visible energy. An excess of events at low hadronic visible energy is observed relative to the prediction of the NEUT event generator. Good agreement is observed with the cross section prediction of the GENIE generator. A search for photons from  $\pi^0$  decay, which would veto a neutral-current  $K^+$  event in a proton decay search, is performed, and a  $2\sigma$  deficit of detached photons is observed relative to the GENIE prediction.

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Proton decay is predicted by grand unification theories [1–13]. Models that incorporate supersymmetry predict the dominant channel to be  $p \to K^+ \bar{\nu}$  [6–13]. Some models

[8-13] predict proton lifetimes greater than  $10^{34}$  years, consistent with the current 90% confidence experimental bound of  $5.6 \times 10^{33}$  years by Super-Kamiokande [14].

In Super-K [15], the  $K^+$  from  $p \to K^+ \bar{\nu}$  is below the Cherenkov threshold of  $T_K = 252$  MeV, where  $T_K$  is the  $K^+$  kinetic energy. Events are selected by reconstructing the  $K^+$  decay products using three techniques as described in Ref. [14]. In the analysis with the lowest predicted background rate, a photon from the deexcitation of the residual nucleus is required, followed in time by a  $\mu^+$  from the decay  $K^+ \to \mu^+ \nu_{\mu}$ , then a "Michel" electron from  $\mu^+ \to e^+ \nu_e \bar{\nu}_{\mu}$ , and no other particles. Neutral-current (NC) production of  $K^+$  by atmospheric neutrinos produces the same signature when no final-state particles are above the Cherenkov threshold. For example,  $\nu O^{16} \to \nu K^+ \Lambda N^{15}$ followed by  $\Lambda \to p\pi^-$  and  $K^+ \to \mu^+ \nu_{\mu}$  is indistinguishable from proton decay when the outgoing  $K^+$ , proton, and  $\pi^$ are below threshold, and the  $\pi^-$  captures.

With 40% photocoverage, the Super-K background prediction based on the Honda atmospheric flux [16] and the NEUT [17] neutrino interaction generator is 1.1 events per Mton yr, of which NC  $K^+$  production is the largest single source at 48% of the total background. In a 260 kiloton-year exposure, Super-K observed zero events with a predicted background of 0.38 events [14]. The proposed Hyper-K [18] will have an exposure of several Mton yrs, in which several background events from NC  $K^+$  production by atmospheric neutrinos would be expected. Constraining this background with a measurement of the cross section for such a process is an important input to future proton decay searches.

DUNE also plans to search for  $p \rightarrow K^+ \bar{\nu}$  [19]. With very low thresholds for charged hadrons, DUNE can potentially reach lower background rates by detecting the associated antikaon or hyperon present in all NC  $K^+$  reactions. However, low-energy hadrons could be present in proton decay signal events due to interactions between the  $K^+$  and the residual nucleus. The event selection must tolerate some nuclear activity to achieve high efficiency, and NC  $K^+$ production with very low visible energy can mimic the signal process. The predicted background rate comes from a single event in a 1 Mton yr simulation with a  $K^+$ , no other tracks or  $\pi^0$ s, and less than 800 MeV of total visible energy [20].

Strange particle production by neutrinos has previously been observed in bubble chambers [21-31] and in NOMAD [32]. Gargamelle observed three NC  $K^+$  production events [31], and reported a neutral- to charged-current ratio for strangeness production.

In this Letter, we report measurements of NC  $K^+$  production by neutrinos in MINERvA with a sample of 201 events after background subtraction. We report differential cross sections with respect to the  $K^+$  kinetic energy, and with respect to a measure of the non- $K^+$  visible energy. To address potential proton decay backgrounds for water Cherenkov detectors, we also quantify the rate of Michel electrons and detached photons in our data relative to the prediction of GENIE, the event generator used in MINERvA, DUNE, and many other experiments. These measurements

serve as benchmarks to cross section predictions used in background simulations for proton decay experiments.

MINERvA is a dedicated neutrino-nucleus cross section experiment in the NuMI beam line [33] at Fermilab. The data used in this Letter were taken in a  $\nu_{\mu}$ -enriched beam with a peak neutrino energy of 3.5 GeV between March 2010 and April 2012, corresponding to  $3.51 \times 10^{20}$  protons on target. A GEANT4-based model is used to simulate the neutrino beam, and is described in Refs. [34,35]. This model is constrained to reproduce thin-target hadron production measurements on carbon [36–40].

The MINERvA detector consists of a core of scintillator strips surrounded by electromagnetic and hadronic calorimeters. For this result, the interaction vertex is constrained to be within the central 108 planes of the scintillator tracking region and no closer than 22 cm to any edge of the planes. The fiducial volume is 5.57 metric tons, consisting of 95% CH and 5% other materials by mass. The design, calibration, and performance of the MINERvA detector are described in Ref. [41].

Neutrino interactions are simulated using the GENIE 2.8.4 neutrino event generator [42]. Strange particles are produced primarily from hadronization of deep inelastic scattering events, and also in the decays of  $\Delta$  and N baryon resonances, for example,  $\Delta(1750) \rightarrow K\Sigma$ . Inelastic reactions for final-state hadronic invariant mass W <1.7 GeV are simulated with a tuned model of discrete baryon resonance production [43]. Strange baryon resonances are not simulated. The transition to deep inelastic scattering is simulated using the Bodek-Yang model [44]. Hadronization at higher energies is simulated with the AGKY model [45] based on the gradual transition from Koba-Nielson-Olesen (KNO) scaling [46] to the Lund string model of PYTHIA (version 6) [47] with increasing W. Parameters that control the rate of strange particle production in hadronization are tuned such that rates of  $\Lambda$ and  $K_{S}^{0}$  production on deuterium and neon agree with BEBC [21-24] and Fermilab 15' [25,26] bubble chamber measurements. Final-state interactions (FSI) are modeled using the INTRANUKE package [42,48]. FSI for  $K^+$  in GENIE 2.8.4 include elastic and inelastic scattering but not charge exchange. Pion-induced processes that produce kaons, for example,  $\pi^+ n \to K^+ \Lambda$ , are not included, nor is any FSI for  $\Lambda$  and  $\Sigma$  hyperons.

In the NEUT generator used by Super-K, hadronization is simulated using PYTHIA for W > 2 GeV, and an exclusive  $K^+\Lambda$  process is simulated for W < 2 GeV [49].

We define the signal process as a neutral-current reaction with at least one  $K^+$  exiting the nucleus in which the neutrino interaction occurred, with kinetic energy less than 600 MeV. There is no requirement on the neutrino flavor or helicity.

Stopping kaons are selected by reconstructing the timing signature of a  $K^+$  decay at rest. The general method used to identify  $K^+$  mesons in MINERvA is described in Ref. [50].



FIG. 1. A neutral-current  $K^+$  data candidate event in MIN-ERvA shown in the *X*-*Z* view. This event is the 13th neutrino interaction in beam spill 408, of run 2371, subrun 53. The small triangles represent time-stamped energy deposits in individual scintillator strips. A hypothesis for each particle track is given. The  $\mu^+$  is delayed in time relative to the  $K^+$ . The track labeled as a pion is probably from  $\Sigma^- \rightarrow n\pi^-$  with an unobserved neutron; however, it could also be a  $K^-$  track.

For NC events, however, a muon tag for primary vertices is not available. Consequently, the kinetic energy of candidate  $K^+$  in this analysis is required to exceed the 100 MeV tracking threshold. Kaons with kinetic energies greater than 600 MeV are not contained unless they undergo an interaction in the detector. For interacting kaons, the rangebased  $K^+$  energy reconstruction underestimates the true kinetic energy because some of the  $K^+$  kinetic energy is transferred to other particles, typically nucleons. An example NC  $K^+$  candidate from data is shown in Fig. 1.

NC events with no muon at the neutrino interaction point are selected by requiring that no track other than the  $K^+$ candidate traverse more than 250 g/cm<sup>2</sup> of material in MINERvA, where the side and downstream calorimeters are included. The efficiency of this selection for true NC events is 83%, where the inefficiency is due to energetic noninteracting pions and protons. In simulation, 11% of true charged-current (CC) events have muons with range less than 250 g/cm<sup>2</sup>, which corresponds to 500 MeV of kinetic energy.

The distribution of the longest track range excluding the  $K^+$  candidate is shown in Fig. 2. The arrow shows the selection at 250 g/cm<sup>2</sup>. Backgrounds are divided into four categories described in detail below: CC or NC  $K^+$  events with  $T_K > 600$  MeV, CC  $K^+$  events with  $T_K < 600$  MeV, CC or NC events where a  $K^+$  does not exit the struck nucleus but is produced inside the detector, and events where some other particle is misidentified as a  $K^+$ . Backgrounds are constrained using a sideband formed from events with tracks traversing more than 250 g/cm<sup>2</sup> of material. Events with rear-exiting tracks are excluded from the sideband in order to make the composition of the backgrounds in the signal and sideband regions more similar, as well as to exclude events at low inelasticity  $y = (E_{\nu} - E_{\mu})/E_{\nu}$ , since CC backgrounds in the signal



FIG. 2. The range of the longest reconstructed track in the event excluding the  $K^+$ , in g/cm<sup>2</sup>. The arrow shows selected events, with events to the right forming the CC-like sideband. The highest bin includes all events above 850 g/cm<sup>2</sup> and is not normalized by the bin width, and the lowest bin includes events in which the  $K^+$  is the only reconstructed track.

region are at high y. Backgrounds due to beam pileup are constrained by the method described in Ref. [50]. All other backgrounds are constrained together using the CC-rich sideband, and a scale factor of  $0.96 \pm 0.23$  is applied based on the fit. The scale factor uncertainty includes effects that are highly correlated between the signal and sideband region. The resulting uncertainty on the background-subtracted data is 10%.

A  $K^+$  with kinetic energy greater than 600 MeV exits the inner detector unless it interacts hadronically. Because the kaon energy reconstruction is range based, all accepted events in this kinematic regime have poorly reconstructed energies. In both CC and NC scattering, high-energy kaons originate from the same hadronization models in GENIE. We apply an uncertainty to these events equal to the difference between the spectra produced by the PYTHIA and KNO hadronization models. This results in an uncertainty of  $^{+46}_{-11}\%$  relative to the central value, which is a gradual transition from KNO to PYTHIA with increasing *W* between 2.3 and 3.0 GeV. We treat the highenergy shape of the  $K^+$  spectrum as correlated between CC and NC events, and use the CC-rich sideband to constrain NC  $K^+$  production with  $K^+$  kinetic energy greater than 600 MeV.

The dominant background is due to true CC  $K^+$  events with muon kinetic energy less than 500 MeV. The constraint from the CC-like sideband involves extrapolating in neutrino energy,  $E_{\nu}$ , and y. Sideband events with higher muon energy preferentially come from higher  $E_{\nu}$  and lower y than CC events in the signal region. Uncertainties due to the modeling of the y distribution in GENIE, as well as uncertainties on the flux shape, enter the analysis in the background subtraction [51]. Another important background is due to  $K^+$  produced by hadronic interactions outside the struck nucleus, labeled " $K^+$  in detector" in Fig. 2. These interactions include pion reactions like  $\pi^+n \to K^+\Lambda$ , and  $K^0$  charge exchange  $K^0p \to K^+n$ . This background is present in both CCand NC-rich samples. The largest uncertainty on the rate of these backgrounds is due to the cross section for the hadronic processes in the detector, which is correlated between the CC- and NC-like regions. We place an *a priori* uncertainty of 100% on both classes of events, which are constrained in the background fit.

We report the differential cross section with respect to the  $K^+$  kinetic energy,  $T_K$ , as well as with respect to a measure of non- $K^+$  energy,  $E_{\text{vis}}$ , defined as the sum of the kinetic energy of all  $\pi^{\pm}$ ,  $K^-$ , and protons, and the total energy of all photons,  $\pi^0$ , and  $K^0$ . The energy sum  $E_{\text{vis}}$  includes the prompt decay products of  $\Lambda$  and  $\Sigma$  hyperons. The relationship between the observed energy in MINERvA and the total hadronic energy,  $\nu$ , depends on the relative fraction of  $\nu$  carried by different particle species. The MINERvA detector responds differently to hadronic showers induced by  $p/\pi$  and electromagnetic showers induced by  $\pi^0 \rightarrow \gamma\gamma$ . Neutrons are detected only when they scatter inside the detector and produce charged particles.

The flux-integrated differential cross section per nucleon in bin i is

$$\left(\frac{d\sigma}{dX}\right)_{i} = \frac{\sum_{j} U_{ij} (N_{j} - N_{j}^{bg})}{\epsilon_{i} N_{\text{nuc}} \Phi \Delta_{i}},\tag{1}$$

where *j* is the index of a reconstructed bin of variable  $X = T_K$ ,  $E_{vis}$ ,  $U_{ij}$  is the unsmearing matrix,  $N_j$  is the number of selected events,  $N_j^{bg}$  is the predicted number of background events,  $\epsilon_i$  is the selection efficiency for signal events,  $N_{nuc}$  is the number of nucleons in the fiducial volume,  $\Phi$  is the integrated  $\nu_{\mu}$  flux prediction, and  $\Delta_i$  is the width of bin *i*.

After background subtraction, including subtracting the estimate for events with  $K^+$  kinetic energy greater than 600 MeV, there are 201 signal events in data. The overall sample purity is 41.2% when events with  $T_K > 600$  MeV are considered a background. The data are unfolded using a Bayesian procedure with three iterations [52] described in Refs. [50,51]. The overall selection efficiency is 4.1%. The largest contribution to the inefficiency is the requirement that the  $K^+$  be at rest for of order 10 ns, which removes 50% of the signal events but is necessary to obtain reasonable purity.

The differential cross section with respect to the kaon kinetic energy,  $T_K$ , is shown in Fig. 3, along with predictions from GENIE 2.8.4 and NEUT 5.3.6. We observe a flat shape in  $d\sigma/dT_K$  between 100 and 600 MeV, consistent with the GENIE and NEUT predictions. The total (shape-only)  $\chi^2$  with 5 (4) degrees of freedom is 1.80 (1.05)



FIG. 3. The differential cross section as a function of  $K^+$  kinetic energy  $T_K$ , compared to GENIE and NEUT. The inner (outer) error bars represent the statistical (total) uncertainty.

for GENIE and 6.26 (0.93) for NEUT. This excess with respect to NEUT is not statistically significant. Absolute and fractional uncertainties, as well as covariance matrices, are given in Supplemental Material [53].

Events with a single  $K^+$  track and very little other energy are potential backgrounds in  $p \rightarrow K^+ \bar{\nu}$  searches. The differential cross section with respect to nonkaon visible energy,  $E_{\rm vis}$ , is shown in Fig. 4. The total (shapeonly)  $\chi^2$  with 6 (5) degrees of freedom is 7.31 (6.10) for GENIE and 12.13 (3.42) for NEUT. The absolute  $\chi^2$  for NEUT in the lowest two bins is 10.48 with 2 degrees of freedom. Strangeness production in NEUT is insufficient at low W, which is related to the final-state hadronic energy,  $\nu$ , by  $W^2 = M^2 + 2M\nu - Q^2$ , where M is the nucleon mass and  $Q^2$  is the squared four-momentum transfer to the nuclear



FIG. 4. The differential cross section as a function of non- $K^+$  visible energy  $E_{vis}$ , as defined in the text, compared to GENIE and NEUT. The inner (outer) error bars represent the statistical (total) uncertainty.

system.  $E_{vis}$  is approximately equal to  $\nu - E_K$ , where  $E_K$  is the  $K^+$  total energy, and the difference is due to the exclusion of neutron kinetic energy and charged pion rest masses from  $E_{vis}$ . Absolute and fractional uncertainties, as well as covariance matrices, are given in Supplemental Material [53].

NEUT is used by Super-K to predict the atmospheric neutrino background in the search for  $p \rightarrow K^+ \bar{\nu}$  [14]. Neutrino events satisfying the  $p \rightarrow K^+ \bar{\nu}$  selection are expected to be at  $E_{\rm vis} < 0.8$  GeV, where the excess in our data compared to NEUT is greatest. A  $p\pi^-$  from  $\Lambda$  decay can have up to 630 MeV of visible energy with both particles below Cherenkov threshold. However, Super-K is a zero-threshold detector of  $\pi^+$  and  $\pi^0$  because of the observable decay products,  $\pi^+ \rightarrow \mu^+ \rightarrow e^+$  and  $\pi^0 \rightarrow \gamma\gamma$ . In NC  $K^+$  events with  $E_{\rm vis} < 0.8$  GeV, a final-state  $\pi^+$  is

predicted in 25% of events in GENIE compared to 4% in NEUT, while a final-state  $\pi^0$  is predicted in 35% of events in GENIE and 39% in NEUT. These percentages include pions from prompt hyperon decays. In NEUT, NC  $K^+$  production for W < 2 GeV is simulated as a  $K^+\Lambda$  exclusive process without additional mesons [49]. This process produces  $\pi^0$ s from  $\Lambda \to n\pi^0$ , but never produces  $\pi^+$ . In NEUT, the threshold for  $\Sigma$  baryon associated production, or for  $K^+$ associated production with additional pions, is W = 2 GeV. In GENIE, reactions like  $\nu p \rightarrow \nu K^+ \Sigma^+ \pi^$ or  $\nu p \rightarrow \nu K^+ \Sigma^- \pi^+$  turn on at threshold around W = 1.8 GeV. At low W, NEUT is also missing reactions like  $\nu n \rightarrow \nu K^+ \Sigma^-$  as well as kaon-antikaon associated production, which could be important sources of background to  $p \to K^+ \bar{\nu}$  due to the lack of  $\pi^0$  and  $\pi^+$ .

Estimates of  $\pi^+$  and  $\pi^0$  in  $E_{\rm vis} < 0.8$  GeV events in MINERvA data are obtained from searches for Michel electrons, and for detached photons. Michel electrons are identified in MINERvA by searching for activity near the end point of a charged particle track that is delayed in time and consistent in energy with an electron from  $\mu \rightarrow e\nu\nu$ . Two selection criteria are used. The first is less stringent, with a 28% efficiency to find a Michel in an event with a  $\pi^+$ , and a 5.5% fake rate. A second, more stringent, selection requires delayed activity in all three MINERvA plane orientations. This reduces the rate of false positives to 1.4%, and reduces the efficiency to 12%.

Good agreement with GENIE in the Michel electron rate is observed. With the aggressive selection, a Michel is identified in 12.7% of simulated selected events with  $E_{\rm vis} < 0.8$  GeV, compared to  $(11.8 \pm 2.6)\%$  in data. With the more stringent selection, the rate is 3.8% in simulation and  $(4.0 \pm 1.5)\%$  in data. Our data are not directly comparable to NEUT because this comparison uses selected events and reconstructed quantities.

To search for detached photons, a visual scan was performed by three scanners on events with reconstructed  $E_{\rm vis} < 0.8$  GeV in data and simulation. Photons are expected to arise primarily from  $\pi^0 \rightarrow \gamma\gamma$ , with  $\pi^0$  from

 $\Lambda \rightarrow n\pi^0$ , but could also be due to  $\Sigma^0 \rightarrow \Lambda \gamma$ . A photon was identified in 23% of data events compared to 33% of simulated events, giving a ratio of data to simulation of  $0.70^{+0.15}_{-0.13}$  (stat)  $\pm 0.01$  (syst), where the statistical uncertainty is due to the finite scanning samples of 177 events each in data and simulation, and the systematic uncertainty is the disagreement between scanners. We observe an indication of an overprediction of the photon rate in NC  $K^+$  events that is significant at  $2\sigma$ .

In neutral-current scattering, a  $K^+$  is always produced in association with a hyperon or antikaon. While MINERvA does not reconstruct individual hyperons, the pion content of the final state gives some sensitivity to different final states. For example, a  $\pi^0$  was identified in 40% of  $K^+\Lambda$ simulated events, compared to 18% of  $K^+K^-$  events. Our data prefer a reduction in the  $K^+\Lambda$  rate and a corresponding enhancement of  $K^+K^-$ , which would reduce the  $\pi^0$  rate while preserving the total  $K^+$  cross section.

The statistics obtained by MINERvA in this channel are far larger than what could be achieved with a sensitive detector exposed to atmospheric neutrinos. The expected number of NC  $K^+$  interactions with  $E_{\rm vis} < 0.8$  GeV on a CH target is computed using GENIE [42] for the MINERvA flux [34] and the atmospheric neutrino flux at Kamioka [16]. For this signal, the MINERvA low-energy run of  $3.51 \times 10^{20}$  protons on target and a 5.57-ton fiducial volume is equivalent to a 58.4 Mton yr exposure to atmospheric neutrinos, and a 9.6 Mton yr exposure to atmospheric antineutrinos.

In conclusion, we report the first high-statistics measurement of NC  $K^+$  production by neutrinos. The measured cross section is in disagreement with the NEUT prediction at low  $E_{vis}$ , and in good agreement with the GENIE prediction. Searches were performed for Michel electrons and for photon conversions, which would cause an event to be rejected in a proton decay search. We observe good agreement with GENIE in the rate of Michel electrons in low- $E_{vis}$  events, and a deficit of 30% in the number of photons compared to GENIE. Our data indicate that the NEUT cross section for NC  $K^+$  production is too small by a factor of 2 in the kinematic region of most importance to searches for  $p \to K^+ \bar{\nu}$ .

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- H. Georgi and S. L. Glashow, Phys. Rev. Lett. 32, 438 (1974).
- [2] J. R. Ellis, D. V. Nanopoulos, and J. Walker, Phys. Lett. B 550, 99 (2002).
- [3] N. Arkani-Hamed, S. Dimopoulos, G. F. Giudice, and A. Romanino, Nucl. Phys. B709, 3 (2005).
- [4] A. Hebecker and J. March-Russell, Phys. Lett. B 539, 119 (2002).
- [5] I. R. Klebanov and E. Witten, Nucl. Phys. B664, 3 (2003).
- [6] S. Dimopoulos and H. Georgi, Nucl. Phys. B193, 150 (1981).
- [7] N. Sakai and T. Yanagida, Nucl. Phys. B197, 533 (1982).
- [8] P. Nath, A. H. Chamseddine, and R. Arnowitt, Phys. Rev. D 32, 2348 (1985).
- [9] Q. Shafi and Z. Tavartkiladze, Phys. Lett. B **473**, 272 (2000).
- [10] V. Lucas and S. Raby, Phys. Rev. D 55, 6986 (1997).
- [11] J. C. Pati, Int. J. Mod. Phys. A 18, 4135 (2003).
- [12] K. S. Babu, J. C. Pati, and F. Wilczek, Phys. Lett. B 423, 337 (1998).
- [13] M. L. Alciati, F. Feruglio, Y. Lin, and A. Varagnolo, J. High Energy Phys. 03 (2005) 054.
- [14] K. Abe et al., Phys. Rev. D 90, 072005 (2014).
- [15] S. Fukuda *et al.*, Nucl. Instrum. Methods Phys. Res., Sect. A 501, 418 (2003).
- [16] M. Honda, M. S. Athar, T. Kajita, K. Kasahara, and S. Midorikawa, Phys. Rev. D 92, 023004 (2015).
- [17] Y. Hayato, Acta Phys. Pol. B 40, 2477 (2009).
- [18] K. Abe et al. (Hyper-Kamokande LOI), arXiv:1109.3262.
- [19] R. Acciarri et al., arXiv:1512.06148.

- [20] A. Bueno, A. J Melgarejo, S. Navas, Z. Dai, Y. Ge, M. Laffranchi, A. Meregaglia, and A. Rubbia, J. High Energy Phys. 04 (2007) 041.
- [21] G. T. Jones et al., Z. Phys. C 57, 197 (1993).
- [22] D. Allasia et al., Nucl. Phys. B224, 1 (1983).
- [23] P. Bosetti et al., Nucl. Phys. B209, 29 (1982).
- [24] S. Willocq et al., Z. Phys. C 53, 207 (1992).
- [25] N.J. Baker et al., Phys. Rev. D 34, 1251 (1986).
- [26] D. DeProspo et al., Phys. Rev. D 50, 6691 (1994).
- [27] S. J. Barish et al., Phys. Rev. Lett. 33, 1446 (1974).
- [28] W. A. Mann, T. Kafka, M. Derrick, B. Musgrave, R. Ammar, D. Day, and J. Gress, Phys. Rev. D 34, 2545 (1986).
- [29] N. J. Baker, P. L. Connolly, S. A. Kahn, H. G. Kirk, M. J. Murtagh, R. B. Palmer, N. P. Samios, and M. Tanaka, Phys. Rev. D 24, 2779 (1981).
- [30] D. Son et al., Phys. Rev. D 28, 2129 (1983).
- [31] H. Deden et al., Phys. Lett. B 58B, 361 (1975).
- [32] P. Astier et al., Nucl. Phys. B621, 3 (2002).
- [33] P. Adamson *et al.*, Nucl. Instrum. Methods Phys. Res., Sect. A **806**, 279 (2016).
- [34] L. Aliaga et al., Phys. Rev. D 94, 092005 (2016).
- [35] L. Aliaga, Report No. FERMILAB-THESIS-2016-03, 2016.
- [36] C. Alt et al., Eur. Phys. J. C 49, 897 (2007).
- [37] A. V. Lebedev, Report No. FERMILAB-THESIS-2007-76, 2007.
- [38] J. Paley et al., Phys. Rev. D 90, 032001 (2014).
- [39] S. M. Seun, Report No. FERMILAB-THESIS-2007-61, 2007.
- [40] D.S. Barton et al., Phys. Rev. D 27, 2580 (1983).
- [41] L. Aliaga *et al.*, Nucl. Instrum. Methods Phys. Res., Sect. A 743, 130 (2014).
- [42] C. Andreopoulos *et al.*, Nucl. Instrum. Methods Phys. Res., Sect. A **614**, 87 (2010).
- [43] D. Rein and L. M. Sehgal, Ann. Phys. 133, 79 (1981).
- [44] A. Bodek, I. Park, and U.-K. Yang, Nucl. Phys. B, Proc. Suppl. 139, 113 (2005).
- [45] T. Yang, C. Andreopoulos, H. Gallagher, K. Hofmann, and P. Kehayias, Eur. Phys. J. C 63, 1 (2009).
- [46] Z. Koba, H. B. Nielsen, and P. Olesen, Nucl. Phys. B40, 317 (1972).
- [47] T. Sjostrand, S. Mrenna, and P. Skands, J. High Energy Phys. 05 (2006) 026.
- [48] S. A. Dytman and A. S. Meyer, AIP Conf. Proc. 1405, 213 (2011).
- [49] Y. Hayato (private communication).
- [50] C. M. Marshall et al., Phys. Rev. D 94, 012002 (2016).
- [51] C. M. Marshall, Report No. FERMILAB-THESIS-2016-12, 2016.
- [52] G. D'Agostini, Nucl. Instrum. Methods Phys. Res., Sect. A 362, 487 (1995).
- [53] See Supplemental Material at http://link.aps.org/ supplemental/10.1103/PhysRevLett.119.011802 for details about tables containing differential cross sections, uncertainties, and covariance matrices.